A new methodology for physics perturbations

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Introduction

EPS

model uncertainty of a forecast due to:

• Error in Initial (and boundary) conditions

Model error

- uncertainty of physical parameterization tendencies: SPPT
- uncertainty of parameters inside parameterization: **RP, SPP**
- Stochastic nature of some physical processes
 - cellular automaton
 - multicloud model



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Stochastically Perturbated Parameterization Tendencies



spatial, temporal correlation tuned for best skill



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Random Parameter - Stochastically Perturbed Parameterizations

$$\frac{\partial \psi}{\partial t}|_{\text{pert}} = f(\mathbf{X}; \xi|_{\text{pert}})$$

$$\xi|_{pert} = \exp(\Psi)\xi|_{unpert}$$

 $\Psi \sim \mathcal{N}(\mu, \sigma)$



σ determined by expert view spatial, temporal correlation tuned for best skill



- Can the modularity of ALARO be used to characterize model error/uncertainty?
 - modularity: represent 1 physical process with different schemes
 - reference Model (most sophisticated/best results) vs approximate model (simpler/easier/faster)
 - lower bound for model error
- What is the effect on the ensemble when using perturbations based on this characterization
 - can be applied to all physical process with multiple representations (cloud condensation, turbulence, radiation)
 - test case: deep convection at 4 km resolution



- Deep convection at 4 km
 - reference Model
 - approximate Model
- Definition of model error
 - fluxes
 - time-scale
- Period Domain

Deep Convection Parameterization

CP = ALARO - 3MT





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Explicit treatment of deep convection

NCP = ALARO - STRAPRO



RMI

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• Transport:

$$\epsilon_{\psi}^{\text{trans}} = J_{\psi,\text{3MT}}^{c} + J_{\psi,\text{3MT}}^{td} - J_{\psi,\text{STRAPRO}}^{td}$$

$$\psi = q_{\nu}, q_l, q_i, h, u, \nu$$

• Condensation:

$$\epsilon_{\psi}^{\text{cond}} = F_{\psi,\text{3MT}}^{c} + F_{\psi,\text{3MT}}^{st} - F_{\psi,\text{STRAPRO}}^{st}$$

$$\psi = \nu l, \nu i$$

• Evaporation:

$$\epsilon_{\psi}^{\text{cond}} = F_{\psi,\text{3MT}}^{c} + F_{\psi,\text{3MT}}^{st} - F_{\psi,\text{STRAPRO}}^{st}$$

 $\psi = rv, sv$



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Error Source

- error made during 1 time step
- start from identical atmospheric state
- first timestep = problem with spin up fluxes
- \rightarrow Let model spin up, then deactivate deep convection scheme





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- Need (a lot of) convection
- No steep orography
 - → Indian Ocean
 - \rightarrow Period of enhanced convective activity (active equatorial waves) 1 - 10 April 2009
- forced by ERA5 analysis





100 105 110 115

100 105 110 115

105 110 115

100

Intermezzo: ALARO in the Tropics





Intermezzo: ALARO in the Tropics

Scores - Precipitation





Intermezzo: ALARO in the Tropics

Scores - Atmospheric fields (250 hPa)





Intermezzo: ALARO in the Tropics

Scores - Atmospheric fields (850 hPa)



— NCP — CP

Deactivation of the deep convection scheme





• Evaporation flux error

Deactivation of the deep convection scheme





• Evaporation flux error



vertical profile

- probability distribution
- intervariable correlation
- spatial autocorrelation
- temporal autocorrelation





- vertical profile
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- temporal autocorrelation





- mean and standard deviation
- probability distribution
- intervariable correlation
- spatial autocorrelation
- temporal autocorrelation





- vertical profile
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How to perturb the tendencies, keeping these correlations?



- vertical profile \checkmark
- probability distribution \checkmark
- \bullet intervariable correlation \checkmark
- spatial autocorrelation
- temporal autocorrelation

How to perturb the tendencies, keeping these correlations?



Perturbation scheme



- \bullet vertical profile \checkmark
- probability distribution \checkmark
- \bullet intervariable correlation \checkmark
- correlation with transport flux
- spatial autocorrelation
- temporal autocorrelation

How to perturb the tendencies, keeping these correlations?

What grid columns enter the database?

Only grid-columns with convective activity \longrightarrow Find a general selection criterion:



Perturbation scheme



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- spatial autocorrelation
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Qv transport flux error @ I20

 $\overline{\sigma_u\omega_u}$



h transport flux error @ I20 20090406 0900 UTC +12h



Perturbation scheme



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When and where do we add perturbations?

Indications for convective activity:

- resolved vertical wind (vertically averaged) (OMEGA)
- moisture convergence (MOCON)

Simple YES/NO threshold \rightarrow Tuned together with $\overline{\sigma_u \omega_u}$



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- spatial autocorrelation (\checkmark)
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For every grid point at every time step:

When and where do we add perturba- if (MOCON/OMEGA > THRESHOLD) then: tions?

- sample from database
- add profiles from sample to physical fluxes



same domain during 11 - 20 April 2009

- flux perturbations only (ERA5)
- combined with IC and BC perturbations (ERA5 EDA)

Verification w.r.t ERA5



RMSE w.r.t. ERA5 HRES analysis: 10 member ensemble mean vs 3MT control vs STRAPRO control





RMSE w.r.t. ERA5 HRES analysis: 10 member ensemble mean vs 3MT control vs STRAPRO control





IC and BC perturbations from ERA5 EDA Continuously Ranked Probability Score (CRPS) = Brier score over all thresholds





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RMI

Results

Spread



RMI

Results

Spread





Results - Combination with IC and BC perturbations

Precipitation - CRPS w.r.t. TRMM (Satellite)





Results - Combination with IC and BC perturbations





Conclusion

- New methodology for quantifying (lower bound on) model error
- Applied to deep convection (in tropics)
- Developed proto-type stochastic perturbation scheme
 - Takes into account PDF and vertical and Test-case over Europe multivariate correlations
 - Conserves total water, heat and momentum.
- largest positive impact in upper air
- small impact on precipitation

Outlook

- Find a way to quantify condensationevaporation flux errors
- Take into account temporal autocorrelation
- Replace database by fitted distributions

